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IDENTIFICATION AND CLASSIFICATION OF COMPACT RADIO
SOURCES IN THE M₁₇ REGION WITH VLA OBSERVATIONS

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*IDENTIFICATION AND CLASSIFICATION OF COMPACT RADIO SOURCES IN THE
M17 REGION WITH VLA OBSERVATIONS*

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*For all the children who one day dreamed of being astronomers and admired the stars from the
distance of society*

RESUMEN

Esta tesis realiza un catálogo de radio fuentes compactas (RFC) no sesgado en la región de M17 con sus propiedades y analiza su naturaleza. Los datos fueron tomados usando el VLA actualizado en su configuración A y banda X (8 – 12 GHz), $\sim 0.2''$ de resolución, en tres épocas: 3,4 y 8 de Mayo del 2018. La calibración y autocalibración fueron aplicadas a los datos usando [CASA](#) para finalmente obtener un mapa de $5 \mu\text{Jy beam}^{-1}$ de nivel de ruido. Luego, las RFC fueron extraídas del mapa final usando tres software: Pybdsf, Aegean y BLOBCAT, obteniendo que este último daba los mejores resultados. Así obtuvimos un catálogo preliminar de 200 RFC. Después de esto, obtuvimos bajo inspección visual un total de 12 fuentes extendidas y un catálogo de 182 fuentes compactas. Para inspeccionar más a fondo su naturaleza, buscamos algunas de sus propiedades, como: contrapartidas en otras longitudes de onda, índice espectral, alta variabilidad rápida y polarización. Finalmente, su naturaleza fue discutida en dos grupos principales: fuentes térmicas y no térmicas, obteniendo un total de 26 fuentes térmicas y 63 no térmicas. Esto indica una predominancia de estrellas inesperadas de Clase III ($\sim 35\%$) dentro de la región, concluyendo que M17 contiene un cúmulo evolucionado. Por otro lado, la inspección de su distribución espacial deja en evidencia que no ha ocurrido ningún brote de formación estelar recientemente, y que ésta ha sido gradual a lo largo del tiempo de vida de la nube.

ABSTRACT

This thesis makes a compact radio sources (CRS) catalog not biased in the M17 region with their properties and analyzes their nature. The data were taken using the upgraded VLA in its A configuration and X band (8 – 12 GHz), $\sim 0.2''$ of resolution, in three epochs: May 3, 4 and 8 of 2018. The calibration and self-calibration were applied to the data using CASA to finally obtain a map of $5 \mu\text{Jy beam}^{-1}$ of rms noise level. Then, the CRS were extracted from the final map using three software: Pybdsf, Aegean and BLOBCAT, obtaining that the latter give us the best results. Thus, we obtained a preliminary catalog of 200 CRS. After that, we obtained under a visual inspection a total of 12 extended sources and a catalog of 182 CRS. To further inspect their nature, we looked for some of their properties like: counterparts at other wavelengths, spectral index, rapid high variability and polarization. Finally, their nature were discussed in two main groups: thermal and non-thermal sources, obtaining a total of 26 thermal and 63 non-thermal CRS. This indicates a predominance of unexpected Class III stars ($\sim 35\%$) within the region, concluding that M17 is a evolved cluster. On the other hand, the inspection of their spatial distribution show that no outbreak of star formation has occurred recently, and this has been gradual throughout the lifetime of the cloud.

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